Surface Rupture in Kinematic Ruptures Models for Hayward Fault Scenario Earthquakes

Brad Aagaard



May 21, 2009

Surface Rupture from Kinematic Rupture Models

Holistic approach to calculating surface slip

• Objective

- Calculate surface slip from kinematic rupture models for large scenario earthquakes on the Hayward fault
- Use Monte Carlo technique to turn assemble probability density function for surface slip from deterministic rupture models

• Constraints

- Spectral content of radiated seismic waves match observations
- Reduce coseismic slip in regions with aseismic creep
- Explicitly include constraints on surface slip (in progress)
- Include estimates of afterslip (in progress)



Deterministic Kinematic Rupture Models

Complete description of spatial and temporal evolution of slip

- Spatial description of slip
 - Background slip distribution (if imposing rupture dimensions)
 - Spatial variation of slip (slip heterogeneity)
- Temporal evolution of slip
 - Hypocenter
 - Rupture speed
 - Slip time function



Rupture Lengths and Epicenters

We often impose a rupture length and epicenter in scenarios





Background Slip Distribution

Imposed rupture dimensions control basic parameters

• Calculate Mw using magnitude-area relation (Hanks and Bakun)

$$M_{\rm w} = \begin{cases} 3.98 + \log A & \text{if } A < 468 \text{km}^2 \\ 3.09 + 4/3 \log A & \text{if } A \ge 468 \text{km}^2 \end{cases}$$

• Calculate average slip from rupture area and magnitude

$$D_{avg} = \frac{M_o}{A\mu}$$

• Taper slip along buried edges of rupture



Example of Background Slip Distribution

Hayward South + North, Mw 7.05





Spatial Variation in Slip

Wavenumber spectra characterized by source inversions

• Somerville et al., 1999

$$P(k_x, k_y) = \frac{1}{1 + (a_x^2 k_x^2 + a_y^2 k_y^2)^2}$$
$$\log a_x = -1.72 + 1/2M_w$$
$$\log a_y = -1.93 + 1/2M_w$$

• Mai and Beroza, 2002

$$P(k_x, k_y) = \frac{a_x a_y}{(1 + a_x^2 k_x^2 + a_y^2 k_y^2)^{H+1}}, H \approx 0.75$$
$$\log a_x = -2.5 + 1/2M_w$$
$$\log a_y = -1.5 + 1/3M_w$$



Blending Background and Spatial Variation

Select wavenumber for cross-over using NGA GMPEs









Accounting for Creep: Slip-predictable Approach

Reduce background slip according to creep

- Use model from Funning et al. (2007) to define rate and distribution of creep
- Assume no afterslip (for the moment)

Locked patches

Slip = Elapsed time * Slip rate

• Creeping patches

Slip = Elapsed time * (Slip rate - Creep rate)



Creep on Hayward Fault: Funning et al. Model

Rate and distribution of creep constrained by geodetic data





Slip-predictable Approach: Background Slip







Accounting for Creep: Slip-gradient Approach

Impose gradient in slip in creeping areas

- Impose vertical gradient in background slip in creeping areas delineated by Funning et al. (2007)
- Compute Mw using Hanks and Bakun (2002) magnitude-area relation using effective rupture area (following WG02)



Accounting for Creep: Slip-gradient Approach

Impose gradient in slip in creeping areas

- Impose vertical gradient in background slip in creeping areas delineated by Funning et al. (2007)
- Compute Mw using Hanks and Bakun (2002) magnitude-area relation using effective rupture area (following WG02)

• Features

- Slip more likely to reach surface in larger events
- More of the creeping areas rupture coseismically in larger events
- Assume no afterslip (for the moment)



Slip-gradient Approach: Background Slip







Hayward South + North Rupture Model

Slip distribution with 1 s rupture time contours

(NW) San Pablo Bay

San Jose (SE)





Constraining Surface Slip

Wells and Coppersmith, 1994

• Maximum slip (strike-slip)

 $\log D_{max} = -7.03 + 1.03 M_{\rm w}$

• Average slip (strike-slip)

$$\log D_{avg} = -6.32 + 0.90 M_{\rm w}$$

• Empirical relation for coseismic slip without creep or afterslip



Surface Slip: Accounting for creep and afterslip

• Effect of creep on coseismic rupture

- Reduced magnitude for given rupture area (WG02 effective area)
- Use reduced magnitude in slip/magnitude relation
- Incorporating afterslip (some assumptions)
 - Afterslip releases stress concentrations following rupture
 - Near-surface slip driven by greater slip at depth



Coseismic Slip from Scenario Rupture Models

Fine-tune stochastic parameters and blending for better match



Hayward South, Mw 6.76





Probabilistic Estimation of Surface Offset

Use Monte Carlo approach to calculate probability density function

- Vary magnitude and rupture end points
 - Gutenberg-Richter frequency-magnitude relation
 - Characteristic events
 - Selection of particular scenarios (this study)
- Random seed for stochastic variation of slip
- Vary slip-gradient in accounting for creep
- Hypocenter, slip time function, rupture speed (temporal evolution only)



Long-Term Research Issues

Incorporate more physics into rupture models

Spontaneous dynamic rupture models include fault constitutive behavior (but are currently poorly constrained)

- Do fault constitutive model parameters vary in the near surface?
 - Stable sliding at shallow depths instead of unstable sliding
 - Variation in friction (cohesion and overburden pressure)
- Does coseismic slip decrease at shallow depths?
- Does surface slip have the same spatial variation as deep slip?
- Is surface coseismic slip slower/faster than slip at depth?
- How does creep affect slip?
 - Does effect of creep vary with earthquake magnitude?
 - Correlation of afterslip with creep?

